



**AB  
GRAD  
CON23**

1  
00:00:00,000 --> 00:00:04,220  
hello

2  
00:00:04,230 --> 00:00:13,430  
[Music]

3  
00:00:18,710 --> 00:00:15,589  
pleasure to be here

4  
00:00:20,390 --> 00:00:18,720  
uh so I'm in the program in Applied

5  
00:00:23,450 --> 00:00:20,400  
Mathematics at the University of Arizona

6  
00:00:27,529 --> 00:00:23,460  
but I work with Professor Regis farrier

7  
00:00:30,170 --> 00:00:27,539  
who's an ecologist and a postdoc antenna

8  
00:00:33,530 --> 00:00:30,180  
affolder who is also an ecologist and

9  
00:00:36,889 --> 00:00:33,540  
planetary scientist so we developed

10  
00:00:39,709 --> 00:00:36,899  
these Global models of Europa and so

11  
00:00:41,270 --> 00:00:39,719  
what you see right here is the Europa

12  
00:00:43,670 --> 00:00:41,280  
Clipper Mission which is going to launch

13  
00:00:45,889 --> 00:00:43,680

next year in October hopefully and I

14

00:00:49,250 --> 00:00:45,899

believe it's the largest in terms of

15

00:00:52,190 --> 00:00:49,260

size spacecraft that NASA's ever built

16

00:00:54,049 --> 00:00:52,200

because of those massive solar panels

17

00:00:58,069 --> 00:00:54,059

um because we're when we're five times

18

00:01:00,410 --> 00:00:58,079

away from the Sun at Jupiter I think the

19

00:01:07,250 --> 00:01:00,420

brightness scales like inverse Square

20

00:01:13,070 --> 00:01:09,950

so why do we care about Europa

21

00:01:15,230 --> 00:01:13,080

so beneath the ice crust we have an

22

00:01:17,270 --> 00:01:15,240

ocean which potentially contains two to

23

00:01:20,030 --> 00:01:17,280

three times the amount of water as on

24

00:01:22,429 --> 00:01:20,040

Earth which is when I first discovered

25

00:01:23,810 --> 00:01:22,439

that fact I was just so blown away by

26

00:01:25,910 --> 00:01:23,820

that

27

00:01:29,090 --> 00:01:25,920

um and we have an energy source from

28

00:01:31,310 --> 00:01:29,100

tidal heating uh namely we may have

29

00:01:33,710 --> 00:01:31,320

hydrothermal venting as a result of that

30

00:01:36,230 --> 00:01:33,720

we saw that on Enceladus because we saw

31

00:01:37,910 --> 00:01:36,240

silicates in the plumes we haven't

32

00:01:40,429 --> 00:01:37,920

confirmed that yet in Europa but it's

33

00:01:42,130 --> 00:01:40,439

very likely based on the amount of tidal

34

00:01:46,609 --> 00:01:42,140

energy that we expect

35

00:01:49,789 --> 00:01:46,619

and we also have a rock ocean interface

36

00:01:52,190 --> 00:01:49,799

so that's very essential for you know

37

00:01:55,190 --> 00:01:52,200

cycling chemicals needed for life

38

00:01:57,230 --> 00:01:55,200

uh and we have a lot of convection as

39

00:01:59,450 --> 00:01:57,240

Sarah talked about

40

00:02:02,569 --> 00:01:59,460

um which is very essential also for

41

00:02:07,429 --> 00:02:05,630

so as a result we need these models to

42

00:02:10,130 --> 00:02:07,439

constrain what we're going to see from

43

00:02:12,290 --> 00:02:10,140

the mission so we had juice also the esa

44

00:02:13,190 --> 00:02:12,300

Mission let's not forget about that

45

00:02:15,170 --> 00:02:13,200

um

46

00:02:16,309 --> 00:02:15,180

so we need to model all these different

47

00:02:16,970 --> 00:02:16,319

layers

48

00:02:19,250 --> 00:02:16,980

um

49

00:02:22,250 --> 00:02:19,260

so working off what we did for install

50

00:02:24,290 --> 00:02:22,260

it is to analyze the Cassini data we're

51  
00:02:28,790 --> 00:02:24,300  
going to do that here specifically

52  
00:02:30,229 --> 00:02:28,800  
focusing on the ecosystem aspect

53  
00:02:32,570 --> 00:02:30,239  
um and so what we want to do with that

54  
00:02:33,830 --> 00:02:32,580  
is provide clear constraints on what

55  
00:02:35,750 --> 00:02:33,840  
kind of BIOS Industries we're going to

56  
00:02:39,949 --> 00:02:35,760  
expect based on a number of different

57  
00:02:42,770 --> 00:02:39,959  
hypotheses uh of abiotic

58  
00:02:48,410 --> 00:02:42,780  
possibilities biotic possibilities and

59  
00:02:53,809 --> 00:02:50,690  
so as I said our approach is very

60  
00:02:55,910 --> 00:02:53,819  
multi-level uh we are going to first

61  
00:02:59,390 --> 00:02:55,920  
look at the bottom layer which is the

62  
00:03:00,949 --> 00:02:59,400  
mixing of the core and the ocean to

63  
00:03:02,449 --> 00:03:00,959

assess you know the chemical structure

64

00:03:05,270 --> 00:03:02,459

coming out of you know coming out

65

00:03:07,009 --> 00:03:05,280

through through the water passing

66

00:03:08,869 --> 00:03:07,019

through the core

67

00:03:10,670 --> 00:03:08,879

and what chemicals are going to come

68

00:03:12,589 --> 00:03:10,680

through that we're going to have to

69

00:03:15,229 --> 00:03:12,599

assess what how that mixes in the water

70

00:03:17,509 --> 00:03:15,239

to assess how much uh chemicals we're

71

00:03:20,149 --> 00:03:17,519

going to see coming out of the plumes

72

00:03:22,910 --> 00:03:20,159

and then we also need to take some model

73

00:03:25,330 --> 00:03:22,920

organisms from Earth to understand the

74

00:03:28,130 --> 00:03:25,340

metabolisms that we might expect

75

00:03:30,830 --> 00:03:28,140

along with the ecological and possible

76

00:03:33,229 --> 00:03:30,840

evolutionary dynamics of the life that

77

00:03:35,509 --> 00:03:33,239

we're modeling

78

00:03:38,030 --> 00:03:35,519

so with this framework we're going to

79

00:03:40,369 --> 00:03:38,040

estimate the likelihoods of certain data

80

00:03:44,330 --> 00:03:40,379

sets that we're going to see via a

81

00:03:49,190 --> 00:03:46,490

so to intimidate you a little bit we're

82

00:03:51,050 --> 00:03:49,200

going to have some math

83

00:03:53,530 --> 00:03:51,060

um so let's talk about the mixing we

84

00:03:56,210 --> 00:03:53,540

look at the dimensionless mixing ratio  $X$

85

00:03:58,190 --> 00:03:56,220

and the temperature  $T$  so we have a

86

00:04:00,470 --> 00:03:58,200

partial differential equation which with

87

00:04:02,149 --> 00:04:00,480

spatial and time variables

88

00:04:04,970 --> 00:04:02,159

solving this partial differential

89

00:04:06,710 --> 00:04:04,980

equation gives us this steady state

90

00:04:09,530 --> 00:04:06,720

composition

91

00:04:12,649 --> 00:04:09,540

and temperature of the mixing layer with

92

00:04:16,550 --> 00:04:12,659

the mass flux density

93

00:04:21,650 --> 00:04:18,949

we can compute what we're going to see

94

00:04:25,310 --> 00:04:21,660

coming out of the plumes by integrating

95

00:04:27,890 --> 00:04:25,320

over the entire space so we we integrate

96

00:04:30,650 --> 00:04:27,900

the integrand contains this Mass flux

97

00:04:32,570 --> 00:04:30,660

density and the concentration of a

98

00:04:35,510 --> 00:04:32,580

specific chemical component that we're

99

00:04:39,530 --> 00:04:35,520

looking at so whether that be a glycine

100

00:04:40,790 --> 00:04:39,540

or other amino acids or you know cells

101  
00:04:44,570 --> 00:04:40,800  
potentially

102  
00:04:47,090 --> 00:04:44,580  
so the concentration of this compound at

103  
00:04:49,909 --> 00:04:47,100  
the base of the ocean plume assuming

104  
00:04:51,530 --> 00:04:49,919  
that the buoyant mixing layer mixes

105  
00:04:53,150 --> 00:04:51,540  
together as we described from the

106  
00:04:55,850 --> 00:04:53,160  
previous model

107  
00:04:59,150 --> 00:04:55,860  
is given by this quotient of these

108  
00:05:03,890 --> 00:05:01,129  
so then we want to look at the cell

109  
00:05:05,810 --> 00:05:03,900  
metabolism and output and there's a very

110  
00:05:08,930 --> 00:05:05,820  
large variety of possibilities we could

111  
00:05:10,790 --> 00:05:08,940  
be working with here so we kind of run

112  
00:05:14,450 --> 00:05:10,800  
the model for tons and tons of

113  
00:05:15,890 --> 00:05:14,460

possibilities of Earth analogs

114

00:05:18,770 --> 00:05:15,900

um depending on these environmental

115

00:05:21,830 --> 00:05:18,780

conditions so we model this catabolic

116

00:05:25,189 --> 00:05:21,840

reaction rate of a specific organism

117

00:05:27,110 --> 00:05:25,199

using the Arrhenius law I know that came

118

00:05:29,029 --> 00:05:27,120

up yesterday Arrhenius in one of the

119

00:05:29,770 --> 00:05:29,039

trivia questions

120

00:05:31,790 --> 00:05:29,780

um

121

00:05:33,969 --> 00:05:31,800

uh so

122

00:05:38,390 --> 00:05:33,979

and that's so the Arrhenius law is

123

00:05:40,850 --> 00:05:38,400

specifically this quotient here and we

124

00:05:44,570 --> 00:05:40,860

use the catabolic energy constant for

125

00:05:47,390 --> 00:05:44,580

given by this term and it goes in here

126  
00:05:52,430 --> 00:05:47,400  
to understand the ratio of inactivated

127  
00:05:54,890 --> 00:05:52,440  
to activated enzymes for the reaction

128  
00:05:57,770 --> 00:05:54,900  
now critically we assess the ecological

129  
00:06:00,170 --> 00:05:57,780  
dynamics of the system so

130  
00:06:02,090 --> 00:06:00,180  
in this case we are dealing with one

131  
00:06:06,529 --> 00:06:02,100  
population

132  
00:06:11,210 --> 00:06:07,969  
ode model

133  
00:06:12,610 --> 00:06:11,220  
describing the Dynamics but we've been

134  
00:06:15,170 --> 00:06:12,620  
thinking about incorporating

135  
00:06:17,029 --> 00:06:15,180  
multi-species into this because there's

136  
00:06:18,529 --> 00:06:17,039  
actually a significant dynamical

137  
00:06:21,890 --> 00:06:18,539  
difference when you have competition

138  
00:06:23,689 --> 00:06:21,900

between species in the environment and

139

00:06:26,930 --> 00:06:23,699

that would seriously affect what we'll

140

00:06:29,350 --> 00:06:26,940

see when we go there so we have to we

141

00:06:32,150 --> 00:06:29,360

have to have multiple different

142

00:06:33,950 --> 00:06:32,160

ecosystem Dynamics models

143

00:06:37,510 --> 00:06:33,960

and so we solve this and get a steady

144

00:06:41,870 --> 00:06:37,520

state concentration of a specific uh

145

00:06:46,790 --> 00:06:43,909

so what can we learn from specifically

146

00:06:48,290 --> 00:06:46,800

the Europa Clipper Mission so the Europa

147

00:06:50,870 --> 00:06:48,300

Clipper Mission has some very

148

00:06:53,689 --> 00:06:50,880

interesting instruments we have a

149

00:06:55,969 --> 00:06:53,699

mapping Imaging spectrometer so it will

150

00:06:58,309 --> 00:06:55,979

map the surface and conduct the

151  
00:07:00,469 --> 00:06:58,319  
spectroscopy of the surface and the

152  
00:07:02,510 --> 00:07:00,479  
distribution of the Organics so as

153  
00:07:05,029 --> 00:07:02,520  
things come out of the plumes they they

154  
00:07:06,890 --> 00:07:05,039  
fall onto the surface and we should

155  
00:07:08,390 --> 00:07:06,900  
expect to see an interesting chemical

156  
00:07:11,450 --> 00:07:08,400  
composition I know they're already doing

157  
00:07:14,090 --> 00:07:11,460  
this with jwst from the from a distance

158  
00:07:15,770 --> 00:07:14,100  
they haven't analyzed that data yet but

159  
00:07:19,809 --> 00:07:15,780  
uh should be interesting

160  
00:07:23,330 --> 00:07:19,819  
and then to sort of sniff the exosphere

161  
00:07:26,510 --> 00:07:23,340  
of Europa we have a mass spectrometer

162  
00:07:28,969 --> 00:07:26,520  
Mass specs it'll determine the the

163  
00:07:31,309 --> 00:07:28,979

composition of the of the ocean beneath

164

00:07:33,529 --> 00:07:31,319

by sniffing that exosphere

165

00:07:36,050 --> 00:07:33,539

and then we have a surface dust analyzer

166

00:07:40,670 --> 00:07:36,060

uh it'll have it'll measure like the

167

00:07:44,350 --> 00:07:43,070

and so in this image you can see kind of

168

00:07:50,870 --> 00:07:44,360

a remote

169

00:07:55,909 --> 00:07:53,330

so let's focus on this slide a little

170

00:07:58,909 --> 00:07:55,919

bit so this is the overall framework so

171

00:08:00,650 --> 00:07:58,919

at the top we have those mechanistic

172

00:08:03,290 --> 00:08:00,660

models for the geochemistry and the

173

00:08:05,570 --> 00:08:03,300

biology that we discussed earlier

174

00:08:09,230 --> 00:08:05,580

and so the flow of information goes from

175

00:08:10,909 --> 00:08:09,240

the top as prior probability probability

176

00:08:14,930 --> 00:08:10,919

densities

177

00:08:16,189 --> 00:08:14,940

that flow into as parameters into the

178

00:08:19,129 --> 00:08:16,199

general model

179

00:08:21,409 --> 00:08:19,139

then we pass those parameters into the

180

00:08:22,969 --> 00:08:21,419

biological model for physiology and

181

00:08:24,950 --> 00:08:22,979

evolution

182

00:08:27,529 --> 00:08:24,960

which then goes to our habitability

183

00:08:30,469 --> 00:08:27,539

Criterion so the habitability Criterion

184

00:08:32,269 --> 00:08:30,479

that we generally use here is it's a

185

00:08:33,889 --> 00:08:32,279

differential equation describing the

186

00:08:36,110 --> 00:08:33,899

population so essentially if I were to

187

00:08:38,329 --> 00:08:36,120

take an Earth-like organism drop it onto

188

00:08:40,130 --> 00:08:38,339

Europa would the change in the

189

00:08:42,170 --> 00:08:40,140

population be positive or would it be

190

00:08:44,269 --> 00:08:42,180

zero so the derivative of population

191

00:08:46,970 --> 00:08:44,279

growth essentially

192

00:08:49,730 --> 00:08:46,980

and so we pass that into Bayes theorem

193

00:08:53,630 --> 00:08:49,740

as prior distributions the posterior

194

00:08:57,769 --> 00:08:53,640

distributions uh come from these uh so

195

00:09:00,590 --> 00:08:57,779

we okay actually to be more specific

196

00:09:02,389 --> 00:09:00,600

we simulate all of these models with

197

00:09:05,090 --> 00:09:02,399

thousands of data points

198

00:09:06,829 --> 00:09:05,100

to get what we call pseudo data in

199

00:09:09,769 --> 00:09:06,839

what's called an approximate Bayesian

200

00:09:12,470 --> 00:09:09,779

computation and so we pass that into the

201  
00:09:15,050 --> 00:09:12,480  
model and we take the observational data

202  
00:09:17,090 --> 00:09:15,060  
and we put that into base theorem so if

203  
00:09:19,550 --> 00:09:17,100  
any of you are aware of Bayes theorem

204  
00:09:21,710 --> 00:09:19,560  
that is a way of assessing the

205  
00:09:23,389 --> 00:09:21,720  
probability of a certain event given

206  
00:09:25,310 --> 00:09:23,399  
prior information and given posterior

207  
00:09:27,290 --> 00:09:25,320  
information

208  
00:09:29,090 --> 00:09:27,300  
um and so now we have this cycle of

209  
00:09:30,530 --> 00:09:29,100  
information so after we get a

210  
00:09:32,269 --> 00:09:30,540  
probability from base theorem we can

211  
00:09:34,130 --> 00:09:32,279  
pass it back into the geophysical

212  
00:09:36,350 --> 00:09:34,140  
parameters model and then continue to

213  
00:09:38,410 --> 00:09:36,360

constrain as we get more observational

214

00:09:41,150 --> 00:09:38,420

data

215

00:09:45,050 --> 00:09:41,160

so here's an example of how we did that

216

00:09:47,810 --> 00:09:45,060

with Enceladus for Cassini

217

00:09:50,269 --> 00:09:47,820

so in this graph here you can see the

218

00:09:51,710 --> 00:09:50,279

blue dots are simulations for

219

00:09:53,630 --> 00:09:51,720

uninhabitable

220

00:09:56,110 --> 00:09:53,640

parameter space

221

00:09:59,389 --> 00:09:56,120

the orange dots are

222

00:10:01,970 --> 00:09:59,399

uninhabited but habitable

223

00:10:03,470 --> 00:10:01,980

and the green dots are habitable and

224

00:10:06,110 --> 00:10:03,480

habited

225

00:10:08,090 --> 00:10:06,120

so what you can see uh from analyzing

226

00:10:10,190 --> 00:10:08,100

the methane content of the plume and

227

00:10:13,009 --> 00:10:10,200

Cassini using this framework was that

228

00:10:15,290 --> 00:10:13,019

the vast majority of simulations for all

229

00:10:19,070 --> 00:10:15,300

the possibilities we assessed

230

00:10:20,150 --> 00:10:19,080

fall under the biotic category

231

00:10:24,350 --> 00:10:20,160

um

232

00:10:30,230 --> 00:10:27,889

I mean that speaks for itself right

233

00:10:32,630 --> 00:10:30,240

so um but there's a very important

234

00:10:34,730 --> 00:10:32,640

constraint on this namely the

235

00:10:37,490 --> 00:10:34,740

probability of abiogenesis

236

00:10:39,110 --> 00:10:37,500

so what we do in this model is assess

237

00:10:42,310 --> 00:10:39,120

all the possibilities from probability

238

00:10:45,410 --> 00:10:42,320

zero to probability one of abiogenesis

239

00:10:49,250 --> 00:10:45,420

and so what we found was at a pretty low

240

00:10:52,069 --> 00:10:49,260

probability of abiogenesis uh the from

241

00:10:53,569 --> 00:10:52,079

the previous slide

242

00:10:56,150 --> 00:10:53,579

um

243

00:10:58,490 --> 00:10:56,160

with a low probability of a biogenesis

244

00:11:00,350 --> 00:10:58,500

you see that biotic explanations are the

245

00:11:03,050 --> 00:11:00,360

most probable for what we're seeing the

246

00:11:05,870 --> 00:11:03,060

in the methane content so we really need

247

00:11:08,690 --> 00:11:05,880

to develop our models of abiogenesis to

248

00:11:11,949 --> 00:11:08,700

really limit limit this variable and

249

00:11:16,009 --> 00:11:11,959

constrain this probability

250

00:11:18,769 --> 00:11:16,019

so the conclusion is in order to find

251  
00:11:21,050 --> 00:11:18,779  
life we really have to model these

252  
00:11:22,610 --> 00:11:21,060  
planetary systems with the assumption

253  
00:11:24,590 --> 00:11:22,620  
that there could be life there already

254  
00:11:25,970 --> 00:11:24,600  
because if we don't we're going to show

255  
00:11:28,370 --> 00:11:25,980  
up there and we're going to think okay

256  
00:11:30,530 --> 00:11:28,380  
it's habitable but why because

257  
00:11:32,210 --> 00:11:30,540  
potentially the habitability of a planet

258  
00:11:34,250 --> 00:11:32,220  
is actually coming from the life is

259  
00:11:36,889 --> 00:11:34,260  
already there so we have to incorporate

260  
00:11:39,910 --> 00:11:36,899  
that into our into our potential models

261  
00:11:42,769 --> 00:11:39,920  
so this will allow us to really enhance

262  
00:11:43,449 --> 00:11:42,779  
Mission data analysis

263  
00:11:46,490 --> 00:11:43,459

um

264

00:11:48,889 --> 00:11:46,500

beforehand uh so you know with when we

265

00:11:50,630 --> 00:11:48,899

had Cassini we showed up there and we

266

00:11:53,210 --> 00:11:50,640

had no idea what was going on we're like

267

00:11:55,910 --> 00:11:53,220

wild plumes this is amazing

268

00:11:58,190 --> 00:11:55,920

um but it took many many years to really

269

00:12:00,170 --> 00:11:58,200

assess this so hopefully with Europa

270

00:12:03,350 --> 00:12:00,180

Clipper we have this framework in hand

271

00:12:06,650 --> 00:12:03,360

as a software package so the data can be

272

00:12:09,470 --> 00:12:06,660

fed through real time and the benefit of

273

00:12:11,329 --> 00:12:09,480

this is then we can narrow down in on

274

00:12:13,730 --> 00:12:11,339

specific areas of Europa that are of

275

00:12:16,910 --> 00:12:13,740

interest to us and really maximize the

276

00:12:30,170 --> 00:12:16,920

scientific yield of these missions

277

00:12:37,250 --> 00:12:34,009

uh nice talk uh two questions maybe the

278

00:12:38,630 --> 00:12:37,260

first did you choose a prior for a

279

00:12:41,030 --> 00:12:38,640

biogenesis

280

00:12:42,829 --> 00:12:41,040

yeah so we sampled the entire space from

281

00:12:45,230 --> 00:12:42,839

zero to one probability

282

00:12:47,449 --> 00:12:45,240

in that in those 10 000 or so

283

00:12:49,310 --> 00:12:47,459

simulations

284

00:12:51,710 --> 00:12:49,320

so basically

285

00:12:53,750 --> 00:12:51,720

um we can't pick a prior for abiogenesis

286

00:12:56,150 --> 00:12:53,760

there's no constraint on that variable

287

00:12:57,889 --> 00:12:56,160

uh so essentially we have to run the

288

00:13:00,949 --> 00:12:57,899

model for all possibilities of

289

00:13:03,710 --> 00:13:00,959

abiogenesis so so like I said

290

00:13:06,230 --> 00:13:03,720

specifically with the Enceladus data we

291

00:13:08,449 --> 00:13:06,240

found that uh much less than 50 percent

292

00:13:10,490 --> 00:13:08,459

chance abiogenesis ensures that the

293

00:13:12,650 --> 00:13:10,500

biotic explanation is the most probable

294

00:13:16,009 --> 00:13:12,660

explanation I believe is around 30 or

295

00:13:18,829 --> 00:13:16,019

less probability but if you're beneath

296

00:13:21,350 --> 00:13:18,839

that threshold then really the abiotic

297

00:13:24,350 --> 00:13:21,360

explanations are way more probable even

298

00:13:27,050 --> 00:13:24,360

though the samples from the abiotic

299

00:13:28,670 --> 00:13:27,060

simulations are much smaller but you

300

00:13:30,290 --> 00:13:28,680

know when the probability of evangelists

301

00:13:33,110 --> 00:13:30,300

is really low

302

00:13:34,250 --> 00:13:33,120

it's just not going to happen right so

303

00:13:48,110 --> 00:13:34,260

okay yeah that answered my second

304

00:13:53,810 --> 00:13:51,230

hey thank you for your talk so for

305

00:13:55,250 --> 00:13:53,820

Europa are you also in terms of

306

00:13:58,610 --> 00:13:55,260

metabolism are you thinking about

307

00:14:01,009 --> 00:13:58,620

methanogenesis or yeah okay so that was

308

00:14:03,769 --> 00:14:01,019

the first step to consider um infer

309

00:14:05,569 --> 00:14:03,779

Enceladus it was kind of a the most

310

00:14:07,490 --> 00:14:05,579

obvious thing to look at because of the

311

00:14:11,329 --> 00:14:07,500

methane content

312

00:14:13,670 --> 00:14:11,339

um with Europa it's more complex uh

313

00:14:15,769 --> 00:14:13,680

because we don't have any real data from

314

00:14:17,629 --> 00:14:15,779

Europa and also the environments on

315

00:14:19,670 --> 00:14:17,639

Europa are much more varied you have

316

00:14:21,710 --> 00:14:19,680

subsurface pockets of water along with

317

00:14:23,750 --> 00:14:21,720

the ocean and maybe you have like

318

00:14:26,810 --> 00:14:23,760

biofilms on the bottom of the eye so

319

00:14:29,750 --> 00:14:26,820

it's it's a much more complex thing

320

00:14:31,250 --> 00:14:29,760

um so my hope is to uh run all the

321

00:14:34,490 --> 00:14:31,260

simulations for all the possible

322

00:14:36,650 --> 00:14:34,500

hypotheses that exist currently okay and

323

00:14:39,050 --> 00:14:36,660

I have one little follow-up question so

324

00:14:40,310 --> 00:14:39,060

you very briefly talked about Dynamic

325

00:14:42,290 --> 00:14:40,320

communities and you mentioned

326

00:14:44,930 --> 00:14:42,300

competition are you just looking at

327

00:14:47,389 --> 00:14:44,940

competition or are you looking at other

328

00:14:49,189 --> 00:14:47,399

sorts of consortiums like in a biofilm

329

00:14:53,329 --> 00:14:49,199

there's competition but there's also

330

00:14:55,670 --> 00:14:53,339

lots of sharing going on right so um if

331

00:14:57,530 --> 00:14:55,680

we're including multi-species in our

332

00:15:00,170 --> 00:14:57,540

models we're hoping to incorporate all

333

00:15:01,370 --> 00:15:00,180

the possible interactions between these

334

00:15:02,030 --> 00:15:01,380

species

335

00:15:05,750 --> 00:15:02,040

um

336

00:15:08,449 --> 00:15:05,760

which is really critical I want to re re

337

00:15:10,730 --> 00:15:08,459

uh iterate because you know the

338

00:15:13,370 --> 00:15:10,740

competition can kill off like all of the

339

00:15:16,009 --> 00:15:13,380

life you know so we need to sort of have

340

00:15:18,050 --> 00:15:16,019

an evolutionary perspective also and not

341

00:15:21,230 --> 00:15:18,060

just like a real-time ecological

342

00:15:23,509 --> 00:15:21,240

perspective to assess these things

343

00:15:26,050 --> 00:15:23,519

um so really good question

344

00:15:28,790 --> 00:15:26,060

thank you

345

00:15:31,069 --> 00:15:28,800

very quick questions

346

00:15:32,689 --> 00:15:31,079

hi Emily bear Stanford University I was

347

00:15:35,090 --> 00:15:32,699

wondering so in the biological part of

348

00:15:36,710 --> 00:15:35,100

the model what kind of data goes into

349

00:15:39,290 --> 00:15:36,720

that and how can environmental

350

00:15:40,670 --> 00:15:39,300

microbiologists help modelers improve

351  
00:15:43,730 --> 00:15:40,680  
their models what kind of data do you

352  
00:15:46,490 --> 00:15:43,740  
need for that right so we need Earth

353  
00:15:48,710 --> 00:15:46,500  
analogs primarily we need the you know

354  
00:15:52,009 --> 00:15:48,720  
catabolic reaction parameters

355  
00:15:53,689 --> 00:15:52,019  
for specific Earth analog so you know we

356  
00:15:55,329 --> 00:15:53,699  
might want to go to Antarctica and see

357  
00:15:58,129 --> 00:15:55,339  
what's going on there

358  
00:16:00,290 --> 00:15:58,139  
and that's our best that's our best way

359  
00:16:02,629 --> 00:16:00,300  
to constrain it at the moment

360  
00:16:06,769 --> 00:16:02,639  
so that's how I think biologists really

361  
00:16:09,769 --> 00:16:06,779  
should really help us in that way

362  
00:16:11,269 --> 00:16:09,779  
yeah exactly